

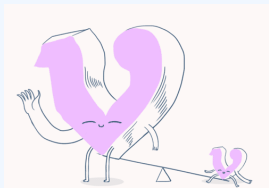
NEUTRINO MASS

ELFT WINTER SCHOOL

TIMO J. KÄRKKÄINEN

EÖTVÖS LORÁND UNIVERSITY

3.2.21



CONTENTS OF TODAY'S PRESENTATION

- First half: the essentials

CONTENTS OF TODAY'S PRESENTATION

- First half: the essentials
 - ▶ Why neutrino masses are important

CONTENTS OF TODAY'S PRESENTATION

- First half: the essentials
 - ▶ Why neutrino masses are important
 - ▶ Helicity, chirality and handedness

CONTENTS OF TODAY'S PRESENTATION

- First half: the essentials
 - ▶ Why neutrino masses are important
 - ▶ Helicity, chirality and handedness
 - ▶ Dirac and Majorana fermions

CONTENTS OF TODAY'S PRESENTATION

- First half: the essentials
 - ▶ Why neutrino masses are important
 - ▶ Helicity, chirality and handedness
 - ▶ Dirac and Majorana fermions
 - ▶ Dirac and Majorana mass terms

- First half: the essentials
 - ▶ Why neutrino masses are important
 - ▶ Helicity, chirality and handedness
 - ▶ Dirac and Majorana fermions
 - ▶ Dirac and Majorana mass terms
 - ▶ Seesaw mechanism, Type I

CONTENTS OF TODAY'S PRESENTATION

- First half: the essentials
 - ▶ Why neutrino masses are important
 - ▶ Helicity, chirality and handedness
 - ▶ Dirac and Majorana fermions
 - ▶ Dirac and Majorana mass terms
 - ▶ Seesaw mechanism, Type I
 - ▶ MeV scale sterile neutrinos
- Second half: alternatives to standard approach

CONTENTS OF TODAY'S PRESENTATION

- First half: the essentials
 - ▶ Why neutrino masses are important
 - ▶ Helicity, chirality and handedness
 - ▶ Dirac and Majorana fermions
 - ▶ Dirac and Majorana mass terms
 - ▶ Seesaw mechanism, Type I
 - ▶ MeV scale sterile neutrinos
- Second half: alternatives to standard approach
 - ▶ Seesaw mechanism, Type II

CONTENTS OF TODAY'S PRESENTATION

- First half: the essentials
 - ▶ Why neutrino masses are important
 - ▶ Helicity, chirality and handedness
 - ▶ Dirac and Majorana fermions
 - ▶ Dirac and Majorana mass terms
 - ▶ Seesaw mechanism, Type I
 - ▶ MeV scale sterile neutrinos
- Second half: alternatives to standard approach
 - ▶ Seesaw mechanism, Type II
 - ▶ Nonstandard interactions (NSI)

CONTENTS OF TODAY'S PRESENTATION

- First half: the essentials
 - ▶ Why neutrino masses are important
 - ▶ Helicity, chirality and handedness
 - ▶ Dirac and Majorana fermions
 - ▶ Dirac and Majorana mass terms
 - ▶ Seesaw mechanism, Type I
 - ▶ MeV scale sterile neutrinos
- Second half: alternatives to standard approach
 - ▶ Seesaw mechanism, Type II
 - ▶ Nonstandard interactions (NSI)
 - ▶ Loop-generated neutrino mass

MOTIVATION: WHY NEUTRINO MASSES ARE IMPORTANT?

- Present bounds from [PLANCK 2018] and [KATRIN 2019]:

$$\sum_{\text{active}} m_\nu \leq 0.12 \text{ eV}, \quad m(\nu_e) < 1.1 \text{ eV}$$

MOTIVATION: WHY NEUTRINO MASSES ARE IMPORTANT?

- Present bounds from [PLANCK 2018] and [KATRIN 2019]:

$$\sum_{\text{active}} m_\nu \leq 0.12 \text{ eV}, \quad m(\nu_e) < 1.1 \text{ eV}$$

- Experimental evidence from **neutrino oscillations**:

$$\Delta m_{21}^2 = (7.42 \pm 0.20) \times 10^{-5} \text{ eV}^2, \quad \Delta m_{32}^2 \stackrel{\text{(NH)}}{=} (2.51 \pm 0.03) \times 10^{-3} \text{ eV}^2$$
$$\Rightarrow \sum_{\text{active}} m_\nu \geq 0.05 \text{ eV}$$

⇒ **Josu Hernández-García's** lecture

MOTIVATION: WHY NEUTRINO MASSES ARE IMPORTANT?

- Present bounds from [PLANCK 2018] and [KATRIN 2019]:

$$\sum_{\text{active}} m_\nu \leq 0.12 \text{ eV}, \quad m(\nu_e) < 1.1 \text{ eV}$$

- Experimental evidence from **neutrino oscillations**:

$$\Delta m_{21}^2 = (7.42 \pm 0.20) \times 10^{-5} \text{ eV}^2, \quad \Delta m_{32}^2 \stackrel{\text{(NH)}}{=} (2.51 \pm 0.03) \times 10^{-3} \text{ eV}^2$$
$$\Rightarrow \sum_{\text{active}} m_\nu \geq 0.05 \text{ eV}$$

⇒ **Josu Hernández-García's** lecture

- ν is the most common **matter** particle in the universe. Its mass gives a small effect on evolution of large-scale structure of the universe:

$$\frac{n_\nu}{n_\gamma} = \frac{3}{11}, \quad \frac{\Omega_\nu}{\Omega_m} \leq 1\%$$

MOTIVATION: WHY NEUTRINO MASSES ARE IMPORTANT?

- Massive neutrinos imply indirectly the existence of massive sterile right-handed neutrinos ν_R (or other new physics...)

MOTIVATION: WHY NEUTRINO MASSES ARE IMPORTANT?

- Massive neutrinos imply indirectly the existence of massive sterile right-handed neutrinos ν_R (or other new physics...)
- Several BSM/GUT's predict massive neutrinos and ν_R : seesaw, left-right symmetry, SO(10)...

$$\mathbf{16} = \begin{pmatrix} \nu_L & d_R^c & d_G^c & d_B^c & u_R & u_G & u_B & e^+ \\ e^- & u_R^c & u_G^c & u_B^c & d_R & d_G & d_B & (\nu_L)^c \end{pmatrix}$$

MOTIVATION: WHY NEUTRINO MASSES ARE IMPORTANT?

- Massive neutrinos imply indirectly the existence of massive sterile right-handed neutrinos ν_R (or other new physics...)
- Several BSM/GUT's predict massive neutrinos and ν_R : seesaw, left-right symmetry, SO(10)...

$$\mathbf{16} = \begin{pmatrix} \nu_L & d_R^c & d_G^c & d_B^c & u_R & u_G & u_B & e^+ \\ e^- & u_R^c & u_G^c & u_B^c & d_R & d_G & d_B & (\nu_L)^c \end{pmatrix}$$

- Sterile ν is a natural DM candidate. \Rightarrow [Károly Sellar's](#) lecture

MOTIVATION: WHY NEUTRINO MASSES ARE IMPORTANT?

- Massive neutrinos imply indirectly the existence of massive sterile right-handed neutrinos ν_R (or other new physics...)
- Several BSM/GUT's predict massive neutrinos and ν_R : seesaw, left-right symmetry, SO(10)...

$$\mathbf{16} = \begin{pmatrix} \nu_L & d_R^c & d_G^c & d_B^c & u_R & u_G & u_B & e^+ \\ e^- & u_R^c & u_G^c & u_B^c & d_R & d_G & d_B & (\nu_L)^c \end{pmatrix}$$

- Sterile ν is a natural DM candidate. \Rightarrow [Károly Sellar's](#) lecture
- Massive neutrinos may take part in CP-violating processes, which are needed to provide baryonic matter-antimatter asymmetry in the universe, either via

MOTIVATION: WHY NEUTRINO MASSES ARE IMPORTANT?

- Massive neutrinos imply indirectly the existence of massive sterile right-handed neutrinos ν_R (or other new physics...)
- Several BSM/GUT's predict massive neutrinos and ν_R : seesaw, left-right symmetry, SO(10)...

$$\mathbf{16} = \begin{pmatrix} \nu_L & d_R^c & d_G^c & d_B^c & u_R & u_G & u_B & e^+ \\ e^- & u_R^c & u_G^c & u_B^c & d_R & d_G & d_B & (\nu_L)^c \end{pmatrix}$$

- Sterile ν is a natural DM candidate. \Rightarrow [Károly Sellar's](#) lecture
- Massive neutrinos may take part in CP-violating processes, which are needed to provide baryonic matter-antimatter asymmetry in the universe, either via
 - ▶ sterile neutrino oscillations in the early universe [[Akhmedov, Rubakov, Smirnov 1998](#)]

MOTIVATION: WHY NEUTRINO MASSES ARE IMPORTANT?

- Massive neutrinos imply indirectly the existence of massive sterile right-handed neutrinos ν_R (or other new physics...)
- Several BSM/GUT's predict massive neutrinos and ν_R : seesaw, left-right symmetry, SO(10)...

$$\mathbf{16} = \begin{pmatrix} \nu_L & d_R^c & d_G^c & d_B^c & u_R & u_G & u_B & e^+ \\ e^- & u_R^c & u_G^c & u_B^c & d_R & d_G & d_B & (\nu_L)^c \end{pmatrix}$$

- Sterile ν is a natural DM candidate. \Rightarrow **Károly Sellar's** lecture
- Massive neutrinos may take part in CP-violating processes, which are needed to provide baryonic matter-antimatter asymmetry in the universe, either via
 - ▶ sterile neutrino oscillations in the early universe [Akhmedov, Rubakov, Smirnov 1998]
 - ▶ leptogenesis [Fukugita, Yanagida 1986] \Rightarrow **Josu Hernández-García's** lecture

- Neutrinos are light, neutral left-chiral spin- $\frac{1}{2}$ leptons.

- Neutrinos are light, neutral left-chiral spin- $\frac{1}{2}$ leptons.
- Spinors (Cartan, 1913) are complex 4D vectors, needed to describe the interactions of spin- $\frac{1}{2}$ particles.

- Neutrinos are light, neutral left-chiral spin- $\frac{1}{2}$ leptons.
- Spinors (Cartan, 1913) are complex 4D vectors, needed to describe the interactions of spin- $\frac{1}{2}$ particles.
- **Dirac spinor** ψ is a solution to **Dirac equation**, and the spinor can be split two chiral components:

$$\psi = \underbrace{\frac{1}{2}(I - \gamma_5)\psi}_{\equiv P_L} + \underbrace{\frac{1}{2}(I + \gamma_5)\psi}_{\equiv P_R} \equiv (P_L + P_R)\psi \equiv \psi_L + \psi_R$$

- Neutrinos are light, neutral left-chiral spin- $\frac{1}{2}$ leptons.
- Spinors (Cartan, 1913) are complex 4D vectors, needed to describe the interactions of spin- $\frac{1}{2}$ particles.
- **Dirac spinor** ψ is a solution to **Dirac equation**, and the spinor can be split two chiral components:

$$\psi = \underbrace{\frac{1}{2}(I - \gamma_5)\psi}_{\equiv P_L} + \underbrace{\frac{1}{2}(I + \gamma_5)\psi}_{\equiv P_R} \equiv (P_L + P_R)\psi \equiv \psi_L + \psi_R$$

- P_L and P_R are **chiral projection operators**, and ψ_L and ψ_R are **eigenstates of chirality**.

- Helicity operator for spin- $\frac{1}{2}$ particles is defined

$$h = \frac{\boldsymbol{\Sigma} \cdot \mathbf{p}}{|\mathbf{p}|}, \quad \boldsymbol{\Sigma} = \gamma^5 \gamma^0 \boldsymbol{\gamma} = \begin{pmatrix} \boldsymbol{\sigma} & 0 \\ 0 & \boldsymbol{\sigma} \end{pmatrix}$$

in Dirac representation of γ -matrices.

- Helicity operator for spin- $\frac{1}{2}$ particles is defined

$$h = \frac{\boldsymbol{\Sigma} \cdot \mathbf{p}}{|\mathbf{p}|}, \quad \boldsymbol{\Sigma} = \gamma^5 \gamma^0 \boldsymbol{\gamma} = \begin{pmatrix} \boldsymbol{\sigma} & 0 \\ 0 & \boldsymbol{\sigma} \end{pmatrix}$$

in Dirac representation of γ -matrices.

- Eigenstates of chirality (ψ_L and ψ_R) coincide for **massless particles** with eigenstates of helicity (ψ_- and ψ_+), corresponding to eigenvalues $\lambda_h = \mp 1$.

- Helicity operator for spin- $\frac{1}{2}$ particles is defined

$$h = \frac{\boldsymbol{\Sigma} \cdot \mathbf{p}}{|\mathbf{p}|}, \quad \boldsymbol{\Sigma} = \gamma^5 \gamma^0 \boldsymbol{\gamma} = \begin{pmatrix} \boldsymbol{\sigma} & 0 \\ 0 & \boldsymbol{\sigma} \end{pmatrix}$$

in Dirac representation of γ -matrices.

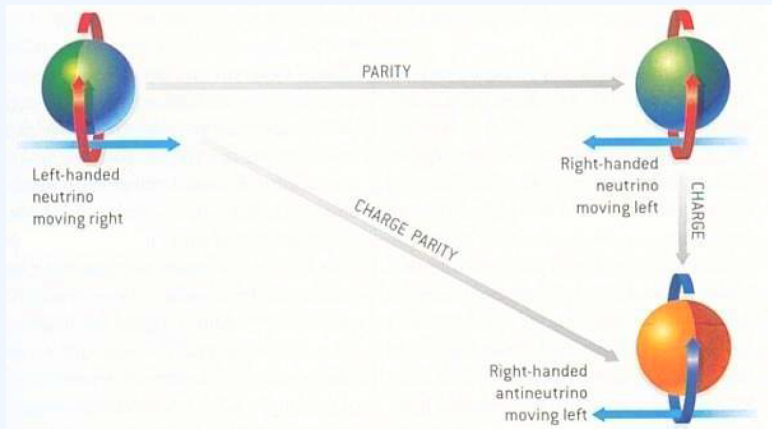
- Eigenstates of chirality (ψ_L and ψ_R) coincide for **massless particles** with eigenstates of helicity (ψ_- and ψ_+), corresponding to eigenvalues $\lambda_h = \mp 1$.
- **In general this is not true!** For low mass limit $m \ll E$

$$\psi_L \approx \psi_- + \frac{m}{E} \psi_+$$
$$\psi_R \approx \psi_+ + \frac{m}{E} \psi_-$$

WEAK INTERACTION IS MAXIMALLY PARITY VIOLATING

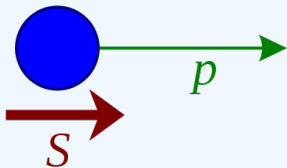
Right-handed (RH) fermion has $\mathbf{p} \uparrow \uparrow \mathbf{S}$.

Left-handed (LH) fermion has $\mathbf{p} \uparrow \downarrow \mathbf{S}$.

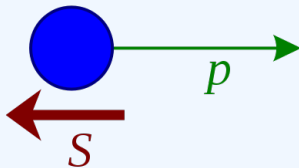


ACTIVE NEUTRINOS CAN BE RIGHT-HANDED

Right-handed:



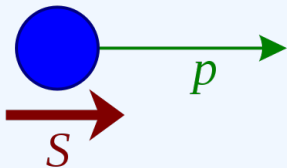
Left-handed:



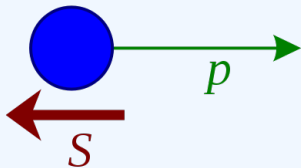
- Since neutrinos are massive, we may Lorentz boost from a frame where neutrino is LH to a frame where it is RH. For massless particles this **helicity flip** is impossible.

ACTIVE NEUTRINOS CAN BE RIGHT-HANDED

Right-handed:



Left-handed:



- Since neutrinos are massive, we may Lorentz boost from a frame where neutrino is LH to a frame where it is RH. For massless particles this **helicity flip** is impossible.
- To write the neutrino mass terms, we introduce six neutrino spinor fields

$$\nu_L \equiv \begin{pmatrix} \nu_{eL} \\ \nu_{\mu L} \\ \nu_{\tau L} \end{pmatrix}, \quad N_R = \begin{pmatrix} N_{1R} \\ N_{2R} \\ N_{3R} \end{pmatrix}$$

- Charge conjugated neutrino fields are defined as

$$(\nu_L)^c \equiv C\bar{\nu}_L^T, \quad (N_R)^c \equiv C\bar{N}_R^T,$$

where $C = i\gamma^2\gamma^0$ is the **charge conjugation operator** in Dirac representation.

- Charge conjugated neutrino fields are defined as

$$(\nu_L)^c \equiv C\bar{\nu}_L^T, \quad (N_R)^c \equiv C\bar{N}_R^T,$$

where $C = i\gamma^2\gamma^0$ is the **charge conjugation operator** in Dirac representation. One can see that

$$\begin{aligned} P_L(N_R)^c &= (N_R)^c, & P_R(\nu_L)^c &= \nu_L^c \\ \Rightarrow (N_R^c) &= (N^c)_L, & (\nu_L)^c &= (\nu^c)_R. \end{aligned}$$

- Charge conjugated neutrino fields are defined as

$$(\nu_L)^c \equiv C\bar{\nu}_L^T, \quad (N_R)^c \equiv C\bar{N}_R^T,$$

where $C = i\gamma^2\gamma^0$ is the **charge conjugation operator** in Dirac representation. One can see that

$$\begin{aligned} P_L(N_R)^c &= (N_R)^c, & P_R(\nu_L)^c &= \nu_L^c \\ \Rightarrow (N_R^c) &= (N^c)_L, & (\nu_L)^c &= (\nu^c)_R. \end{aligned}$$

- Therefore we only need ν_L , N_R and their charge conjugates to write the neutrino mass terms.

$$\mathcal{L} = -\bar{L}_L Y_\nu H' N_R + \text{h.c.}, \quad L_L = \begin{pmatrix} \ell_L \\ \nu_L \end{pmatrix}, \quad H' = i\sigma_2 H^*$$

$$\mathcal{L} = -\bar{L}_L Y_\nu H' N_R + \text{h.c.}, \quad L_L = \begin{pmatrix} \ell_L \\ \nu_L \end{pmatrix}, \quad H' = i\sigma_2 H^*$$

- Weak interaction Lagrangian is written in flavour basis. We choose a basis where the charged lepton masses coincide with their flavour.

$$\mathcal{L} = -\bar{L}_L Y_\nu H' N_R + \text{h.c.}, \quad L_L = \begin{pmatrix} \ell_L \\ \nu_L \end{pmatrix}, \quad H' = i\sigma_2 H^*$$

- Weak interaction Lagrangian is written in flavour basis. We choose a basis where the charged lepton masses coincide with their flavour.
- After spontaneous symmetry breaking $SU(2)_L \otimes U(1)_Y \rightarrow U(1)_Q$,

$$\mathcal{L}' = -\frac{Y_\nu v}{\sqrt{2}} \bar{\nu}_L N_R + \text{h.c.} \equiv -\bar{\nu}_L M_D N_R + \text{h.c.}, \quad v \approx 246.22 \text{ GeV}$$

$$\mathcal{L} = -\bar{L}_L Y_\nu H' N_R + \text{h.c.}, \quad L_L = \begin{pmatrix} \ell_L \\ \nu_L \end{pmatrix}, \quad H' = i\sigma_2 H^*$$

- Weak interaction Lagrangian is written in flavour basis. We choose a basis where the charged lepton masses coincide with their flavour.
- After spontaneous symmetry breaking $SU(2)_L \otimes U(1)_Y \rightarrow U(1)_Q$,

$$\mathcal{L}' = -\frac{Y_\nu v}{\sqrt{2}} \bar{\nu}_L N_R + \text{h.c.} \equiv -\bar{\nu}_L M_D N_R + \text{h.c.}, \quad v \approx 246.22 \text{ GeV}$$

- M_D is the Dirac mass matrix, which can be diagonalized via a biunitary transform: $V^\dagger M_D U = M_{\text{diag}}$.

- In the new basis (**mass basis**),

$$\mathcal{L}' = -\overline{\nu'_L} M_{\text{diag}} \nu'_R + \text{h.c.}, \quad \nu'_L = V^\dagger \nu_L, \quad \nu'_R = U^\dagger \nu_R$$

- In the new basis (**mass basis**),

$$\mathcal{L}' = -\bar{\nu}'_L M_{\text{diag}} N'_R + \text{h.c.}, \quad \nu'_L = V^\dagger \nu_L, \quad N'_R = U^\dagger N_R$$

- Defining a **Dirac spinor** $\nu' = \nu'_L + N'_R = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$,

$$\mathcal{L}' = -\bar{\nu}' M_{\text{diag}} \nu' = -\sum_{i=1}^3 m_i \bar{\nu}_i \nu_i$$

- In the new basis (**mass basis**),

$$\mathcal{L}' = -\bar{\nu}'_L M_{\text{diag}} N'_R + \text{h.c.}, \quad \nu'_L = V^\dagger \nu_L, \quad N'_R = U^\dagger N_R$$

- Defining a **Dirac spinor** $\nu' = \nu'_L + N'_R = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$,

$$\mathcal{L}' = -\bar{\nu}' M_{\text{diag}} \nu' = -\sum_{i=1}^3 m_i \bar{\nu}_i \nu_i$$

- Lepton number is conserved, but lepton flavour violated.

- In the new basis (**mass basis**),

$$\mathcal{L}' = -\bar{\nu}'_L M_{\text{diag}} N'_R + \text{h.c.}, \quad \nu'_L = V^\dagger \nu_L, \quad N'_R = U^\dagger N_R$$

- Defining a **Dirac spinor** $\nu' = \nu'_L + N'_R = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$,

$$\mathcal{L}' = -\bar{\nu}' M_{\text{diag}} \nu' = -\sum_{i=1}^3 m_i \bar{\nu}_i \nu_i$$

- Lepton number is conserved, but lepton flavour violated.
- $\nu\beta\beta$ not possible: $(A, Z) \rightarrow (A, Z + 2) + 2e^-$ violates L .

FERMION MASS TERMS IN THE STANDARD MODEL

$$Y^{\ell} \bar{L}_{L\alpha} H \ell_{R\alpha} + Y^u \bar{Q}_{Li} H u_{Ri} + Y^d \bar{Q}_{Li} H' d_{Ri}$$

FERMION MASS TERMS IN THE STANDARD MODEL

$$Y^{\ell} \bar{L}_{L\alpha} H \ell_{R\alpha} + Y^u \bar{Q}_{Li} H u_{Ri} + Y^d \bar{Q}_{Li} H' d_{Ri}$$

↓ SSB

$$\frac{v}{\sqrt{2}} Y_{\alpha\beta}^{\ell} \ell_{L\alpha} \ell_{R\beta} + \frac{v}{\sqrt{2}} Y_{ij}^u u_{Li} u_{Rj} + \frac{v}{\sqrt{2}} Y_{ij}^d d_{Li} d_{Rj}$$

FERMION MASS TERMS IN THE STANDARD MODEL

$$Y^\ell \bar{L}_{L\alpha} H \ell_{R\alpha} + Y^u \bar{Q}_{Li} H u_{Ri} + Y^d \bar{Q}_{Li} H' d_{Ri}$$

↓ SSB

$$\frac{v}{\sqrt{2}} Y_{\alpha\beta}^\ell \bar{l}_{L\alpha} l_{R\beta} + \frac{v}{\sqrt{2}} Y_{ij}^u \bar{u}_{Li} u_{Rj} + \frac{v}{\sqrt{2}} Y_{ij}^d \bar{d}_{Li} d_{Rj}$$

↓ mass matrices

$$M^\ell = \frac{v}{\sqrt{2}} Y^\ell, \quad M^u = \frac{v}{\sqrt{2}} Y^u, \quad M^d = \frac{v}{\sqrt{2}} Y^d$$

- Yukawa couplings in the SM spans six orders of magnitude without explanation. (Do we **need** explanation?)

$$\begin{array}{lll} Y_e = 3 \times 10^{-6} & Y_\mu = 6 \times 10^{-4} & Y_\tau = 10^{-2} \\ Y_u = 10^{-5} & Y_c = 7 \times 10^{-3} & Y_t = 1 \\ Y_d = 3 \times 10^{-5} & Y_s = 5 \times 10^{-4} & Y_b = 2 \times 10^{-2} \end{array}$$

FLAVOUR PROBLEM GETS WORSE

- Yukawa couplings in the SM spans six orders of magnitude without explanation. (Do we **need** explanation?)

$$\begin{array}{lll} Y_e = 3 \times 10^{-6} & Y_\mu = 6 \times 10^{-4} & Y_\tau = 10^{-2} \\ Y_u = 10^{-5} & Y_c = 7 \times 10^{-3} & Y_t = 1 \\ Y_d = 3 \times 10^{-5} & Y_s = 5 \times 10^{-4} & Y_b = 2 \times 10^{-2} \end{array}$$

- Adding just the Dirac mass terms for neutrinos introduces disturbingly small Yukawas: $Y_\nu \sim \mathcal{O}(10^{-13})$. **Why so small?**

- Yukawa couplings in the SM spans six orders of magnitude without explanation. (Do we **need** explanation?)

$$Y_e = 3 \times 10^{-6}$$

$$Y_\mu = 6 \times 10^{-4}$$

$$Y_\tau = 10^{-2}$$

$$Y_u = 10^{-5}$$

$$Y_c = 7 \times 10^{-3}$$

$$Y_t = 1$$

$$Y_d = 3 \times 10^{-5}$$

$$Y_s = 5 \times 10^{-4}$$

$$Y_b = 2 \times 10^{-2}$$

- Adding just the Dirac mass terms for neutrinos introduces disturbingly small Yukawas: $Y_\nu \sim \mathcal{O}(10^{-13})$. **Why so small?**
- Exotic way out: N_R wave function leaks to extra dimensions?
[Dienes *et al.* 1999, Arkani-Hamed *et al.* 2002]

- All SM fermions are Dirac fermions, but there is an alternative for neutral fermions - neutrinos.

- All SM fermions are Dirac fermions, but there is an alternative for neutral fermions - neutrinos.
- Since $(\nu_L)^c$ is RH, we may write a mass term for neutrinos without postulating the N_R fields:

$$\mathcal{L} = -\frac{1}{2}\overline{\nu_L}M_L(\nu_L)^c + \text{h.c.}$$

NOT-SO-EXOTIC WAY: MAJORANA MASS TERM

- All SM fermions are Dirac fermions, but there is an alternative for neutral fermions - neutrinos.
- Since $(\nu_L)^c$ is RH, we may write a mass term for neutrinos without postulating the N_R fields:

$$\mathcal{L} = -\frac{1}{2}\overline{\nu_L}M_L(\nu_L)^c + \text{h.c.}$$

- This breaks $SU(2)_L \otimes U(1)_Y$ gauge invariance, but we will fix this problem later.

NOT-SO-EXOTIC WAY: MAJORANA MASS TERM

- All SM fermions are Dirac fermions, but there is an alternative for neutral fermions - neutrinos.
- Since $(\nu_L)^c$ is RH, we may write a mass term for neutrinos without postulating the N_R fields:

$$\mathcal{L} = -\frac{1}{2}\bar{\nu}_L M_L (\nu_L)^c + \text{h.c.}$$

- This breaks $SU(2)_L \otimes U(1)_Y$ gauge invariance, but we will fix this problem later.
- A symmetric matrix M_L may be diagonalized via a unitary transformation: $U^T M_L U = M_{\text{diag}}$.

- In the new basis (**mass basis** again)

$$\mathcal{L} = -\frac{1}{2}\overline{\nu'_L} M (\nu'_L)^c + \text{h.c.}, \quad \nu'_L = U^T \nu_L, \quad (\nu'_L)^c \equiv C \overline{\nu'_L}^T$$

MAJORANA MASS TERM

- In the new basis (**mass basis** again)

$$\mathcal{L} = -\frac{1}{2}\overline{\nu'_L} M (\nu'_L)^c + \text{h.c.}, \quad \nu'_L = U^T \nu_L, \quad (\nu'_L)^c \equiv C\overline{\nu'_L}^T$$

- Defining a **Majorana spinor** $\nu' \equiv \nu'_L + (\nu'_L)^c = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$, which satisfies **Majorana condition** $(\nu')^c = \nu'$, that is, neutrinos are their own antiparticles (**Majorana, 1937**).

- In the new basis (**mass basis** again)

$$\mathcal{L} = -\frac{1}{2}\overline{\nu'_L} M (\nu'_L)^c + \text{h.c.}, \quad \nu'_L = U^T \nu_L, \quad (\nu'_L)^c \equiv C \overline{\nu'_L}^T$$

- Defining a **Majorana spinor** $\nu' \equiv \nu'_L + (\nu'_L)^c = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$, which satisfies **Majorana condition** $(\nu')^c = \nu'$, that is, neutrinos are their own antiparticles (**Majorana, 1937**).

$$\mathcal{L} = -\frac{1}{2}\overline{\nu'_L} M_{\text{diag}} (\nu'_L)^c + \text{h.c.} = -\frac{1}{2} \sum_{i=1}^3 m_i \overline{\nu}_i \nu_i$$

MAJORANA MASS TERM

- In the new basis (**mass basis** again)

$$\mathcal{L} = -\frac{1}{2}\overline{\nu'_L} M (\nu'_L)^c + \text{h.c.}, \quad \nu'_L = U^T \nu_L, \quad (\nu'_L)^c \equiv C\overline{\nu'_L}^T$$

- Defining a **Majorana spinor** $\nu' \equiv \nu'_L + (\nu'_L)^c = \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$, which satisfies **Majorana condition** $(\nu')^c = \nu'$, that is, neutrinos are their own antiparticles (**Majorana, 1937**).

$$\mathcal{L} = -\frac{1}{2}\overline{\nu'_L} M_{\text{diag}} (\nu'_L)^c + \text{h.c.} = -\frac{1}{2} \sum_{i=1}^3 m_i \overline{\nu}_i \nu_i$$

- Factor $\frac{1}{2}$ accounts for the fulfillment of Dirac equation for ν_L .

DIRAC VS MAJORANA

In a minimal model, to **only add the neutrino masses** and nothing else:

	Dirac	Majorana
Neutrino oscillations	Yes	Yes
$\nu\nu\beta\beta$	No	Yes
CP violating phases	1	3
Sterile neutrinos	Yes	No
Gauge invariance broken	No	Yes!
Lepton number broken	No	Yes, $\Delta L = 2$
Lepton flavour broken	Yes	Yes
Renormalizable	Yes	Yes

$d = 5$ EFFECTIVE OPERATOR (WEINBERG, 1979)

$$\frac{\mathcal{L}^{(5)}}{\Lambda} = \frac{1}{2} f \overline{L_{L\alpha}} H' (H')^T (L_{L\beta})^c + \text{h.c.}, \quad H' = i\sigma_2 H^*$$

$d = 5$ EFFECTIVE OPERATOR (WEINBERG, 1979)

$$\frac{\mathcal{L}^{(5)}}{\Lambda} = \frac{1}{2} f \overline{L_{L\alpha}} H' (H')^T (L_{L\beta})^c + \text{h.c.}, \quad H' = i\sigma_2 H^*$$

- The only gauge invariant dimension-5 operator which can be constructed from SM fields produces the Majorana mass term after spontaneous symmetry breaking:

$$\frac{\mathcal{L}'}{\Lambda} = -\frac{1}{2} \overline{\nu_L} M_L (\nu_L)^c + \text{h.c.}, \quad M_L = \frac{1}{2} f v^2$$

Λ is the cutoff scale and f neutrino coupling matrix.

$d = 5$ EFFECTIVE OPERATOR (WEINBERG, 1979)

$$\frac{\mathcal{L}^{(5)}}{\Lambda} = \frac{1}{2} f \overline{L_{L\alpha}} H' (H')^T (L_{L\beta})^c + \text{h.c.}, \quad H' = i\sigma_2 H^*$$

- The only gauge invariant dimension-5 operator which can be constructed from SM fields produces the Majorana mass term after spontaneous symmetry breaking:

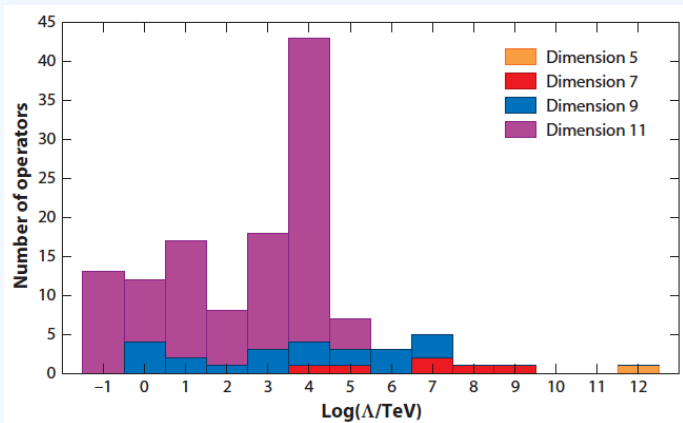
$$\frac{\mathcal{L}'}{\Lambda} = -\frac{1}{2} \overline{\nu_L} M_L (\nu_L)^c + \text{h.c.}, \quad M_L = \frac{1}{2} f v^2$$

Λ is the cutoff scale and f neutrino coupling matrix.

- Higher-dimensional operators exist for neutrino masses, but will not be covered here. Why?

THERE ARE $\mathcal{O}(100)$ DIFFERENT POSSIBILITIES!

Number of possible effective operators increases if higher dimensions are allowed. Assume coupling matrix $f_{ij} = \mathcal{O}(1)$.



Picture: André De Gouvêa, Annu. Rev. Nucl. Part. Sci. 2016. 66:197217.

THEORY IS NOT RENORMALIZABLE!

When we include Weinberg operator...

	Dirac	Majorana
Neutrino oscillations	Yes	Yes
$0\nu\beta\beta$	No	Yes
CP violating phases	1	3
Sterile neutrinos	Yes	No
Gauge invariance broken	No	No
Lepton number broken	No	Yes, $\Delta L = 2$
Lepton flavour broken	Yes	Yes
Renormalizable	Yes	No!

- Type I: add gauge-singlet fermions.

- Type I: add gauge-singlet fermions.
 - ▶ one N_R : only one massive SM neutrino. **No!**

- Type I: add gauge-singlet fermions.
 - ▶ one N_R : only one massive SM neutrino. **No!**
 - ▶ two N_R : minimal seesaw, $m(\nu_1) = 0$. **Possible.**

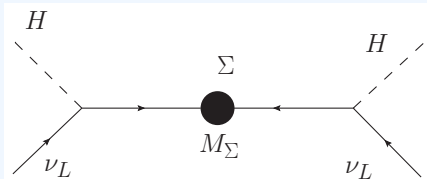
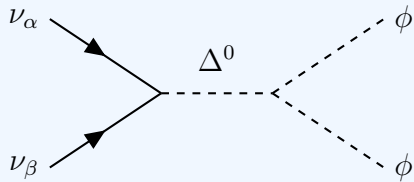
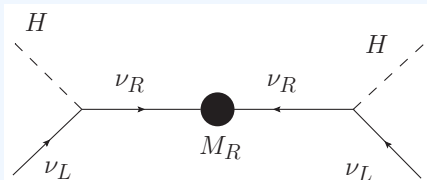
- Type I: add gauge-singlet fermions.
 - ▶ one N_R : only one massive SM neutrino. **No!**
 - ▶ two N_R : minimal seesaw, $m(\nu_1) = 0$. **Possible.**
 - ▶ three N_R : vanilla seesaw. **Most commonly used.**

SEESAW MECHANISMS WORK AT TREE LEVEL

- Type I: add gauge-singlet fermions.
 - ▶ one N_R : only one massive SM neutrino. **No!**
 - ▶ two N_R : minimal seesaw, $m(\nu_1) = 0$. **Possible.**
 - ▶ three N_R : vanilla seesaw. **Most commonly used.**
- Type II: add scalar triplet
 $\Delta = (\Delta^{++}, \Delta^+, \Delta^0) \sim (\mathbf{3}, 2)$

SEESAW MECHANISMS WORK AT TREE LEVEL

- Type I: add gauge-singlet fermions.
 - ▶ one N_R : only one massive SM neutrino. **No!**
 - ▶ two N_R : minimal seesaw, $m(\nu_1) = 0$. **Possible.**
 - ▶ three N_R : vanilla seesaw. **Most commonly used.**
- Type II: add scalar triplet $\Delta = (\Delta^{++}, \Delta^+, \Delta^0) \sim (\mathbf{3}, 2)$
- Type III: add fermion triplet $\Sigma = (\Sigma^+, \Sigma^0, \Sigma^-) \sim (\mathbf{3}, 0)$



TYPE I SEESAW MECHANISM

[Fritzsch, Gell-Mann, Minkowski 1975; Minkowski 1977; Yanagida 1979; Glashow 1979;
Gell-Mann, Ramond, Slansky 1979; Mohapatra and Senjanović 1979

- Use both Dirac and Majorana mass terms and N_R .

$$\mathcal{L} = -\bar{L}_L Y_\nu H' N_R - \underbrace{\frac{1}{2} \bar{N}_R M_R (N_R)^c}_{\text{gauge invariant!}} + \text{h.c.}$$

TYPE I SEESAW MECHANISM

[Fritzsch, Gell-Mann, Minkowski 1975; Minkowski 1977; Yanagida 1979; Glashow 1979; Gell-Mann, Ramond, Slansky 1979; Mohapatra and Senjanović 1979]

- Use both Dirac and Majorana mass terms and N_R .

$$\mathcal{L} = -\bar{L}_L Y_\nu H' N_R - \underbrace{\frac{1}{2} \bar{N}_R M_R (N_R)^c}_{\text{gauge invariant!}} + \text{h.c.}$$

- After SSB, we may construct the full 6×6 mass matrix:

$$\mathcal{L}_m = -\frac{1}{2} \begin{pmatrix} \overline{(\nu_L)^c} & N_R \end{pmatrix} \begin{pmatrix} \mathbf{0} & M_D \\ M_D^T & M_R \end{pmatrix} \begin{pmatrix} \nu_L \\ (N_R)^c \end{pmatrix} + \text{h.c.}$$

TYPE I SEESAW MECHANISM

[Fritzsch, Gell-Mann, Minkowski 1975; Minkowski 1977; Yanagida 1979; Glashow 1979; Gell-Mann, Ramond, Slansky 1979; Mohapatra and Senjanović 1979]

- Use both Dirac and Majorana mass terms and N_R .

$$\mathcal{L} = -\bar{L}_L Y_\nu H' N_R - \underbrace{\frac{1}{2} \overline{N_R} M_R (N_R)^c}_{\text{gauge invariant!}} + \text{h.c.}$$

- After SSB, we may construct the full 6×6 mass matrix:

$$\mathcal{L}_m = -\frac{1}{2} \overline{(\nu_L)^c}, \quad N_R \begin{pmatrix} \mathbf{0} & M_D \\ M_D^T & M_R \end{pmatrix} \begin{pmatrix} \nu_L \\ (N_R)^c \end{pmatrix} + \text{h.c.}$$

- Once we block-diagonalize the mass matrix, we obtain in leading order

$$m_\nu = -M_D M_R^{-1} M_D^T = -\frac{v^2}{2} Y_\nu M_R^{-1} Y_\nu^T, \quad M_N = M_R$$

- If we avoid flavour problem ($Y_{ij} \sim 1$), then we need the RH neutrinos N_R to be extremely heavy: $M_N = \mathcal{O}(10^{15})$ GeV, near GUT scale.
⇒ no hope for direct detection

$$m_\nu = -\frac{v^2}{2} Y_\nu M_R^{-1} Y_\nu^T$$
$$\underset{Y \sim 1}{\approx} 0.01 \text{ eV} \times \frac{10^{15} \text{ GeV}}{M_R}$$

NATURAL TYPE-I SEESAW MECHANISM

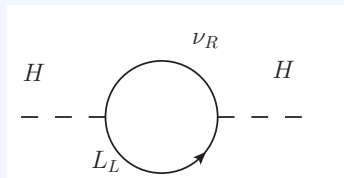
- If we avoid flavour problem ($Y_{ij} \sim 1$), then we need the RH neutrinos N_R to be extremely heavy: $M_N = \mathcal{O}(10^{15})$ GeV, near GUT scale.

⇒ no hope for direct detection

- **Hierarchy problem** gets worse: Higgs mass gets a large one-loop correction.

$$m_\nu = -\frac{v^2}{2} Y_\nu M_R^{-1} Y_\nu^T$$
$$\underset{Y \sim 1}{\approx} 0.01 \text{ eV} \times \frac{10^{15} \text{ GeV}}{M_R}$$

$$\delta M_H^2 = -\frac{\text{eig}(Y_\nu)^2}{8\pi^2} \left(\Lambda^2 + M_{N_i}^2 \ln \frac{M_{N_i}^2}{\Lambda^2} \right)$$



- However the seesaw mechanism works also for lower scales, even though there is **no theoretical justification** for choosing such a scale.

- However the seesaw mechanism works also for lower scales, even though there is **no theoretical justification** for choosing such a scale.
- The mixing between active and sterile neutrinos is characterized by **active-sterile mixing matrix**, which can be obtained from mass matrix block-diagonalization:

$$U_{\ell i} = M_D M_R^{-1} = \frac{v}{\sqrt{2}} Y_\nu M_R^{-1}$$
$$\Rightarrow |U_{\ell i}|^2 = \frac{m_{\nu\ell}}{M_{N_i}} = \mathcal{O}(10^{-11}) \frac{\text{GeV}}{M_{N_i}}$$

- However the seesaw mechanism works also for lower scales, even though there is **no theoretical justification** for choosing such a scale.
- The mixing between active and sterile neutrinos is characterized by **active-sterile mixing matrix**, which can be obtained from mass matrix block-diagonalization:

$$U_{\ell i} = M_D M_R^{-1} = \frac{v}{\sqrt{2}} Y_\nu M_R^{-1}$$
$$\Rightarrow |U_{\ell i}|^2 = \frac{m_{\nu \ell}}{M_{N_i}} = \mathcal{O}(10^{-11}) \frac{\text{GeV}}{M_{N_i}}$$

- This value acts as an **approximate lower bound** for the simplest seesaw case.

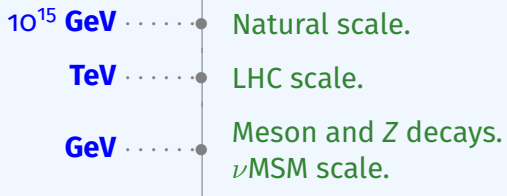
DIFFERENT SEESAW SCALES FOR DIFFERENT PROBLEMS

10^{15} GeV ● Natural scale.

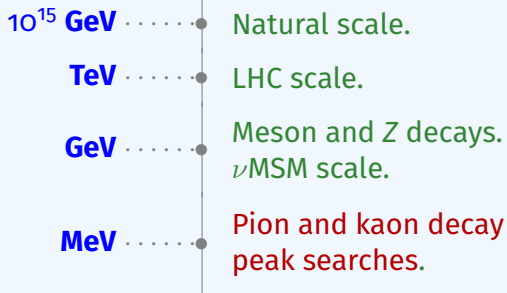
DIFFERENT SEESAW SCALES FOR DIFFERENT PROBLEMS

10^{15} **GeV** ● Natural scale.
TeV ● LHC scale.

DIFFERENT SEESAW SCALES FOR DIFFERENT PROBLEMS



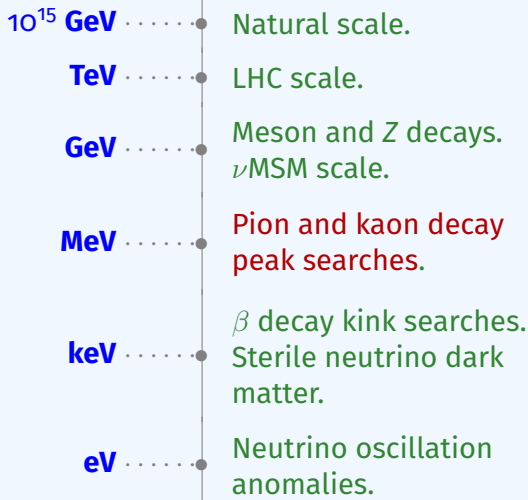
DIFFERENT SEESAW SCALES FOR DIFFERENT PROBLEMS



DIFFERENT SEESAW SCALES FOR DIFFERENT PROBLEMS

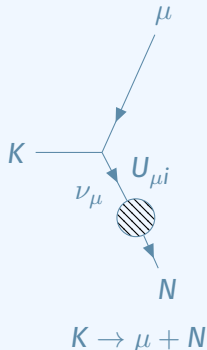
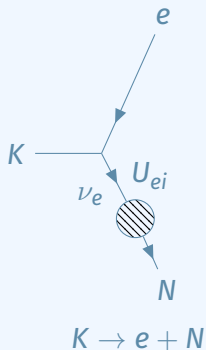
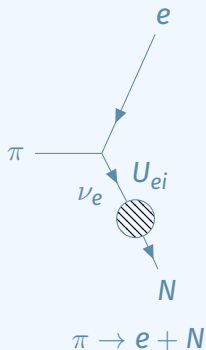


DIFFERENT SEESAW SCALES FOR DIFFERENT PROBLEMS



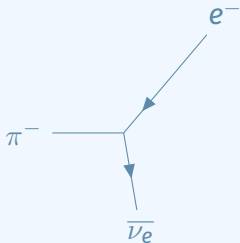
STERILE NEUTRINO PRODUCTION

If $M_{\text{meson}} > M_N + m_e$, sterile neutrino N is produced like an active neutrino, but the **rate is suppressed by $|U_{ei}|^2$** .



EXAMPLE: $\pi^- \rightarrow e^- \bar{\nu}_e$ AT REST

$$\Gamma(\pi^- \rightarrow e^- \bar{\nu}_e) = \frac{G_F^2 f_\pi^2 \cos^2 \theta_c m_\pi^3}{8\pi} \left(\frac{m_e}{m_\pi} \right)^2 \left(1 - \frac{m_e^2}{m_\pi^2} \right)$$

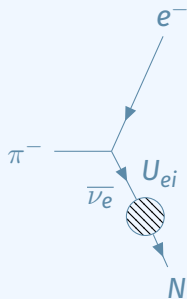


Pion has spin 0, so e^- and $\bar{\nu}_e$ must have opposite spins. At limit $m_\nu \approx 0$, $\bar{\nu}_e$ is purely RH and therefore has positive helicity, so e^- must be LH. Due to conservation of angular momentum, e^- must also have positive helicity, but

$$\psi_L \approx \psi_- + \frac{m_e}{E} \psi_+$$

Helicity suppression.

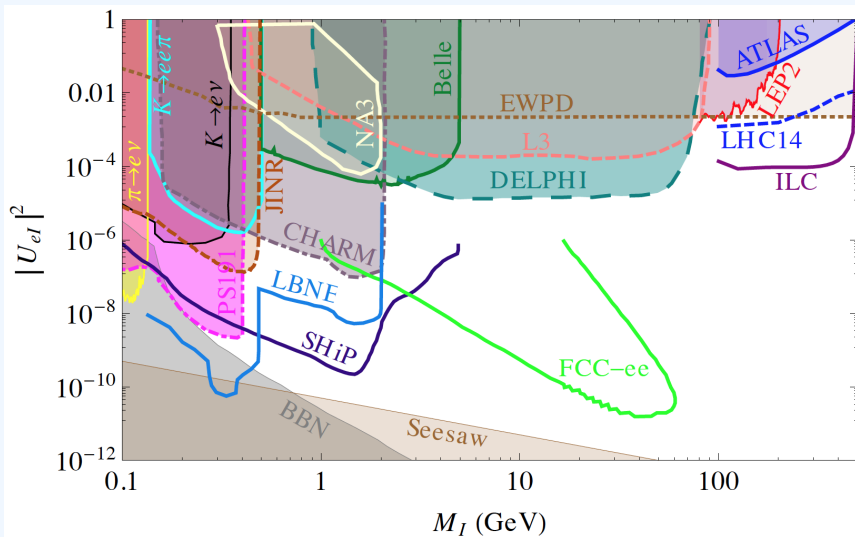
EXAMPLE: $\pi^- \rightarrow e^- N$ AT REST



$$\Gamma(\pi^- \rightarrow e^- N) = \Gamma(\pi^- \rightarrow e^- \bar{\nu}_e) |U_{ei}|^2 \left(\frac{M_N}{m_e}\right)^2$$

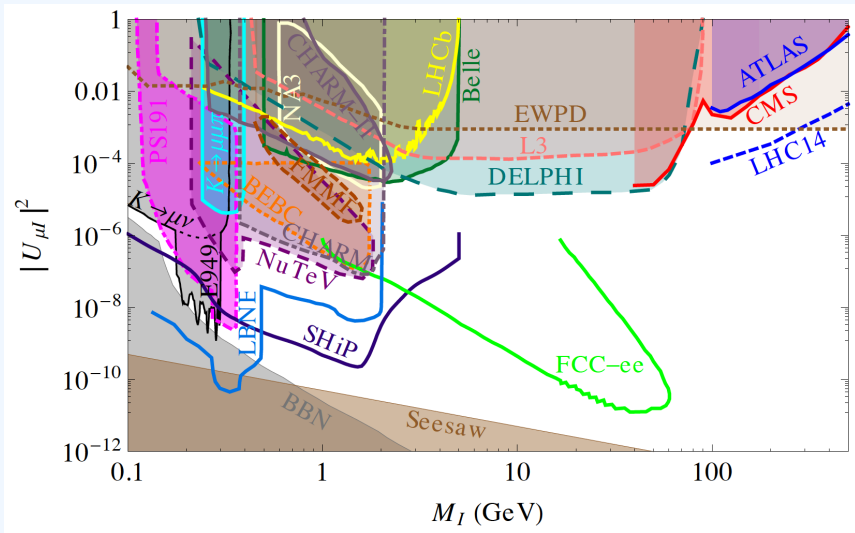
Being massive, sterile neutrino has both helicities. Therefore the electron will not be helicity suppressed, and there will be **a peak** on the pion decay spectrum. Similar searches can be done with kaon decays.

BOUNDS ON ν_e -STERILE MIXING STRENGTH



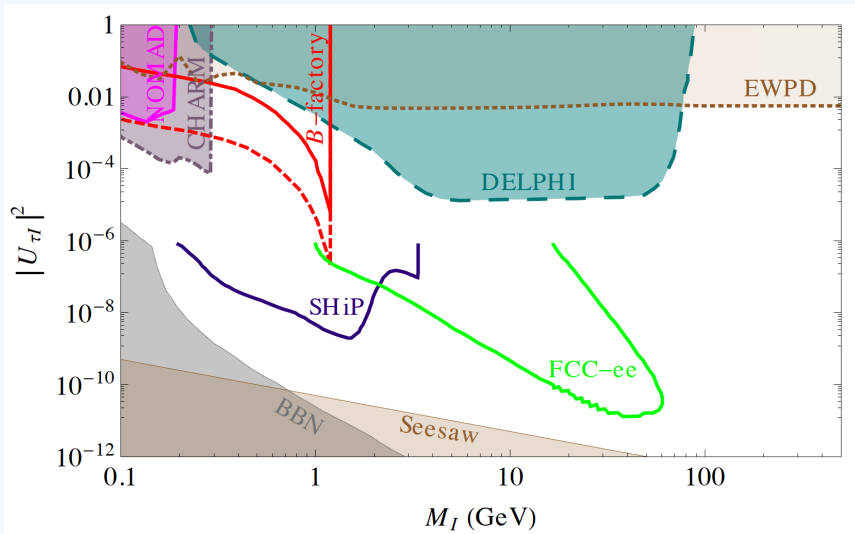
Plot: Alekhin et al., Rep. Prog. Phys. 79 (2016) 124201

BOUNDS ON ν_μ -STERILE MIXING STRENGTH



Plot: Alekhin et al., Rep. Prog. Phys. 79 (2016) 124201

BOUNDS ON ν_τ -STERILE MIXING STRENGTH



Plot: Alekhin et al., Rep. Prog. Phys. 79 (2016) 124201

IN CONCLUSION OF THIS LECTURE

- Neutrinos are weakly interacting, very light, spin- $\frac{1}{2}$, flavour-oscillating elementary particles.

IN CONCLUSION OF THIS LECTURE

- Neutrinos are weakly interacting, very light, spin- $\frac{1}{2}$, flavour-oscillating elementary particles.
- Chirality and helicity coincide for massless particles.

IN CONCLUSION OF THIS LECTURE

- Neutrinos are weakly interacting, very light, spin- $\frac{1}{2}$, flavour-oscillating elementary particles.
- Chirality and helicity coincide for massless particles.
- Dirac and Majorana mass terms can be combined to produce light neutrino masses via seesaw mechanism, which utilizes gauge-singlet heavy sterile Majorana neutrinos.

IN CONCLUSION OF THIS LECTURE

- Neutrinos are weakly interacting, very light, spin- $\frac{1}{2}$, flavour-oscillating elementary particles.
- Chirality and helicity coincide for massless particles.
- Dirac and Majorana mass terms can be combined to produce light neutrino masses via seesaw mechanism, which utilizes gauge-singlet heavy sterile Majorana neutrinos.
- Low-scale seesaw implies sterile neutrinos at mass range available for several different experimental approaches.

THANK YOU FOR YOU ATTENTION!