Search for SUSY at the LHC

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Outline

- Problems of the Standard Model
- Supersymmetry (SUSY): a solution for all?
- Searching for SUSY phenomena
- Natural SUSY: finished?
- Status and prospects

Best introduction to SUSY:

A Supersymmetry Primer by Stephen P. Martin https://arxiv.org/abs/hep-ph/9709356v7 (last revised in 2016)

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SM: parameter fitting, 2018



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SM at CERN's souvenir shop





John Ellis, the creator h.c. = Herm. conj. or hot coffee?

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CMS: 78 p-p collisions in one event!



An event (i.e. bunch crossing) can have many p-p collisions. This increases the statistics and complicates the analysis.

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ATLAS vs. CMS

Both optimized to detect (and study!) the production and decay of Higgs-bosons. Very different detectors giving very similar results.

	ATLAS	CMS	
Magnet	toroid + <i>small(?)</i> 2 T solenoid	large 3.8 T solenoid	
Tracker	semiconductor + TRD	semiconductor	
E-m calorimeter	LAr with steel and Pb	PbWO ₄ scint.	
Hadron calm.	steel + scint. tiles brass + scint. tiles		
Far forward h-cal	LAr with Cu and W	steel with quartz Cher.	
Muon detector	chambers (4 types)	chambers (3 types)	
Size	$\oslash 25\mathrm{m} imes46\mathrm{m}$ (23000 m 3)	$\oslash 15\text{m} imes 21.6\text{m}$ (3800 m ³)	
Trigger	3-level	2-level	
Weight	7000 t	14000 t	
Participants (sci)	3000	2300	

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CMS & ATLAS, signal strengths

So how Higgs study can help to find new physics? Compare precise measurements with SM predictions

Production rate (cross section) ratios: $\mu_i = \sigma_i/(\sigma_i)_{
m SM}$ (*i* = ggF, VBF, WH, ZH, ttH)

Relative decay rates (ratios of branching fractions): $\mu^f = B^f/(B^f)_{\rm SM}$ (f = ZZ, WW, $\gamma\gamma$, $\tau\tau$, bb, $\mu\mu$)

Production and decay cannot be separated, what is really measured:

$$\mu_i^f = rac{\sigma_i B^f}{(\sigma_i)_{
m SM} (B^f)_{
m SM}} = \mu_i \mu^f$$

Allowing for BSM interpretation: $\sigma_i B^f = \sigma_i(\underline{\kappa}) \cdot \Gamma^f(\underline{\kappa}) / \Gamma_H$ Γ_H, Γ^f : total and frac. decay widths, Coupling modifiers: $\kappa_j^2 = \sigma_j / \sigma_j^{SM}$ (prod.); $\kappa_j^2 = \Gamma^j / \Gamma_{SM}^j$ (decay)

[ATLAS and CMS Collaborations], JHEP 1608 (2016) 045.

CMS & ATLAS, signal strengths, Run 1



-p.10/38

Production channels: CMS, $H \rightarrow \gamma \gamma$



ATLAS & CMS: coupling mod's, Run 1



Coupling to fermions & bosons: per expt. and average per decay mode

[ATLAS and CMS Collaborations], JHEP 1608 (2016) 045.

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Problems of the Standard Model – 1

- Gravity? S = 2 graviton?
- Asymmetries: right ⇔ left World ⇔ Antiworld
- Artificial mass creation: Higgs-field ad hoc
- Charge quantization: $Q_{\mathrm{e}} = Q_{\mathrm{p}}, \; Q_{\mathrm{d}} = Q_{\mathrm{e}}/3$
- Why the 3 fermion families?
- Nucleon spin: how 1/2 produced?
- 19 free parameters (too many ??):
 - 3 couplings: α , Θ_W , $\Lambda_{\rm QCD}$; 2 Higgs: M_H , λ
 - 9 fermion masses: $3 \times M_{\ell}, \ 6 \times M_q$
 - 4 parameters of the CKM matrix: Θ_1 , Θ_2 , Θ_3 , δ
 - QCD-vacuum:

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Problems of the Standard Model – 2

Neutrino mysteries

- $M_{\nu} > 0 \Rightarrow +3$ masses, +4 mixing matrix The SM does not like them...
- Are $\nu_{\rm R}$, $\overline{\nu}_{\rm L}$ sterile with no interaction?
- What makes them oscillate (extra interaction)?
- Are they Majorana particles $\overline{\nu} \equiv \nu$?
- Gravitational mass of the Universe:
 - 4% ordinary matter (H, He, stars, gas, dust, ν)
 - 23% invisible dark matter (out of SM!)
 - 73% mysterious dark energy
- Naturalness (hierarchy):

The mass of the Higgs boson quadratically diverges due to radiative corrections. Cancelled if fermions and bosons exist in pairs!

Hierarchy problem

BEH potential: $V = -m_H^2 \Phi^2 + \lambda \Phi^4$ $m_H = 125 \text{GeV}$ Corrections to the mass of the Higgs boson: Fermion loop: f $\Delta m_H^2 = rac{|\lambda_f|^2}{8\pi^2} [-\Lambda_{
m UV}^2 + 6m_f^2 \ln rac{\Lambda_{
m UV}}{m_f} + \ldots]$ Η Η For the top quark $\lambda_t pprox 1$ UV cutoff: $\Lambda_{\rm UV} > 10^{14} {
m GeV!}$ Tuning of the BEH parameters with the precision of 10^{-12} !

If there were two heavy scalar particles:

$$\mathrm{f} \leftrightarrow \mathrm{S}_1, \mathrm{S}_2$$
: $\lambda_S = \lambda_f^2; \;\; m_S = m_f$



 $\Delta m_{H}^{2} = rac{\lambda_{S}}{16\pi^{2}} [+\Lambda_{\mathrm{UV}}^{2} - 2m_{S}^{2} \ln rac{\Lambda_{\mathrm{UV}}}{m_{S}} + \ldots]$ Λ^{2} corrections would cancel out

Question: Why do we need two scalars for one fermion?

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Many-many different models





Beyond the Standard Model



Y. Gershtein et al., "Working Group Report: New Particles, Forces, and Dimensions,"

arXiv:1311.0299.

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Supersymmetry (SUSY)

Hypothesis: Fermions and bosons exist in pairs: $Q|F>=|B>; Q|B>=|F> m_B = m_F$ Identical particles, just spins different

Broken at low energy, partners: much larger mass?



SUSY should solve many problems



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SUSY: Higgs sector

2 Higgs doublets \Rightarrow masses to upper and lower fermions

Extended left-right asymmetry:

 $m_L=m_R$, but $ilde{m}_L
eq ilde{m}_R$

Question: Why should the scalar SUSY partners of the left- and right-polarized SM fermions be distinctly different particles with different masses?

8 Higgs fields \Rightarrow 5 Higgs bosons: h^0, H^0, A^0, H^{\pm}

Higgs parameters: $tan\beta = v_1/v_2$, masses

Question: How can the two Higgs (BEH) doublet fields result in five Higgs-bosons? Why do the 5 Higgs bosons have four higgsino partners (and not five) in the SUSY space?

LSP = dark matter?

SUSY's quantum number: R parity $R = (-1)^{3B-L+2S}$ (B: baryon charge, L: lepton charge, S: spin)

> R = +1 particle, R = -1 SUSY partner Parity-like: $R^2 = +1$

If *R* conserved, lightest SUSY particle (LSP) is stable

R parity may not be much violated: we would detect LSP decays

Neutral LSP: excellent dark matter candidate



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Minimal Supersymmetric SM

Electroweak symmetry breaking \Rightarrow MSSM-fermions mix into mass eigenstates {Electroweak gauginos + higgsinos} \Rightarrow {charginos and neutralinos } $\left\{\tilde{B}(=\tilde{\gamma}), \tilde{W}^{\pm}, \tilde{W}^{0}(=\tilde{Z}); \tilde{h}^{0}, \tilde{H}^{0}, \tilde{H}^{\pm}\}\right\} \Rightarrow \left\{\tilde{\chi}_{1}^{\pm}, \tilde{\chi}_{2}^{\pm}; \tilde{\chi}_{1}^{0}, \tilde{\chi}_{2}^{0}, \tilde{\chi}_{3}^{0}, \tilde{\chi}_{4}^{0}\right\}$

(mass grows with index) Question: gluinos do not mix?

Lightest SUSY particle (LSP) depends on model, e.g. $\tilde{\chi}_1^0$

SUSY breaking (how?) \Rightarrow many (> 100) new parameters masses, couplings, mixing angles Lots of model variants, huge parameter space, different constraints. hMSSM: m(h) = 125 GeV (with corrections!)

Questions: Why should neutralinos be Majorana particles (their own antiparticles)? What should prevent the LSP as dark matter to make dark galaxies and black holes?

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The missing MSSM menagerie

Kind	spin	R parity	gauge eigenstate	mass eigenstate
Higgs bosons	0	+1	${ m H}^0_1, { m H}^0_2, { m H}^+_1, { m H}^2$	$\mathrm{h^0, H^0, A^0, H^\pm}$
			$ ilde{\mathrm{u}}_L, ilde{\mathrm{u}}_R, ilde{\mathrm{d}}_L, ilde{\mathrm{d}}_R$	same
squark	0	-1	$ ilde{ extsf{s}}_L, ilde{ extsf{s}}_R, ilde{ extsf{c}}_L, ilde{ extsf{c}}_R$	same
			$ ilde{\mathrm{t}}_L, ilde{\mathrm{t}}_R, ilde{\mathrm{b}}_L, ilde{\mathrm{b}}_R$	$ ilde{ extbf{t}}_1, ilde{ extbf{t}}_2, ilde{ extbf{b}}_1, ilde{ extbf{b}}_2$
			$ ilde{\mathrm{e}}_L, ilde{\mathrm{e}}_R, ilde{ u}_\mathrm{e}$	same
slepton	0	-1	$ ilde{\mu}_L, ilde{\mu}_R, ilde{ u}_\mu$	same
			$ ilde{ au}_L, ilde{ au}_R, ilde{ u}_{ au}$	$ ilde{ au}_1, ilde{ au}_2, ilde{ u}_{ au}$
neutralino	1/2	-1	$ ilde{\mathrm{B}^0}, ilde{\mathrm{W}^0}, ilde{\mathrm{H}}^0_1, ilde{\mathrm{H}}^0_2$	$ ilde{\chi}^0_1, ilde{\chi}^0_2, ilde{\chi}^0_3, ilde{\chi}^0_4$
chargino	1/2	-1	$ ilde{\mathrm{W}}^{\pm}, ilde{\mathrm{H}}_1^+, ilde{\mathrm{H}}_2^-$	$ ilde{\chi}_1^\pm, ilde{\chi}_2^\pm$
gluino	1/2	-1	ĝ	same
goldstino	1/2	-1	Ĝ	same
gravitino	3/2			

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SUSY: coupling constants



Minimal Supersymmetric Standard Model: Unification! Bend at low energies: SUSY enters with many new particles \Rightarrow more loop corrections

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MSSM mass spectrum: preconceptions

Even if we remain sceptic it is worthwhile to know what do most of the model constructors think (after S.P. Martin)

- R parity is barely violated
- LSP: $ilde{\chi}_1^0$ or gravitino
- ${}$ Gluino mass $M_3\equiv m(ilde g)\gg m(ilde\chi_1^0),m(ilde\chi_2^0),m(ilde\chi_1^\pm)$
- $m(\tilde{u}_i) \sim m(\tilde{d}_i) \sim m(\tilde{c}_i) \sim m(\tilde{s}_i) \gg m(\tilde{\ell}_i)$
- $m(\tilde{u}_i) \sim m(\tilde{d}_i) \sim m(\tilde{c}_i) \sim m(\tilde{s}_i) > (0, 6_{\text{MSUGRA}} \dots 0, 8_{\text{GMSB}}) m(\tilde{g})$
- $m(\tilde{u}_L) \ge m(\tilde{u}_R) \dots m(\tilde{s}_L) \ge m(\tilde{s}_R)$ and $m(\tilde{e}_L) \ge m(\tilde{e}_R), m(\tilde{\mu}_L) \ge m(\tilde{\mu}_R)$ as $M_L^2 \sim M_R^2 + 0, 5m_{1/2}^2$.
- ${old p} \ m({
 m h}^0) \lesssim 150 \ {
 m GeV} \ll m({
 m A}), m({
 m H}^{\pm}), m({
 m H}^0)$

SUSY search

Production in pairs, decay to other SUSY particle (if *R* conserved)

Lightest (LSP) stable, neutral, not observable Signal: missing energy

Hypothetical SUSY decays (LSP = $\tilde{\chi}_1^0$):

- squark: $\tilde{q} \rightarrow q + \tilde{g}; q + \tilde{\chi}_1^0$
- slepton: $\tilde{\ell} \rightarrow \ell + \tilde{\chi}_1^0$
- gluino: $\tilde{g} \rightarrow q + \overline{q} + \tilde{\chi}_1^0$; $g + \tilde{\chi}_1^0$
- wino: $ilde{W}
 ightarrow ext{e} + u_{ ext{e}} + ilde{\chi}_1^0$

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Simplified Models

Few on-shell particles, simple topology and decays Not model-independent, but possibly associated with several models. Possible new physics on well understood SM-base

What can we learn of such analysis?

- Boundaries of search sensitivity, both for data analysis and for new theories.
- Characterizing new physics signals: what models can be associated?
- Limits on more general models: from possible cross-sections.

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Simplified models for SUSY searches

Basic topologies with no lepton and missing energy (signal of Lightest SUSY Particle, LSP):



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Search methods: examples

• α_T search for early discovery in (forced) 2-jet events $(E_T(J_1) > E_T(J_2))$: Cut $\alpha_T = \frac{E_T(J_2)}{M_T(J_1,J_2)}$ $- E_T(J_2)$

 $= \frac{1}{\sqrt{(E_T(J_1) + E_T(J_2))^2 - (p_x(J_1) + p_x(J_2))^2 - (p_y(J_1) + p_y(J_2))^2}}$ Exclusive 2-jet, inclusive 3-jet search

- Jets + H_T for > 2 jets, inclusive Scalar mom. sum: $H_T = \sum_i |\underline{p}_T(J_i)|;$ Missing transverse mom.: $MHT = H_T = |-\sum_i \underline{p}_T(J_i)|$
- Razor search: test kinematic consistency for pair production of heavy particles Two jets (inv. mass M_R) + 0 or 1 lepton

Experimental limits, constraints

No SUSY phenomenon observed, the data limit the parameter space

- LEP, Tevatron, LHC: Higgs sector
 - Mass of SM Higgs from direct searches $M_{
 m H} = 125~{
 m GeV};~~{
 m H} \sim {
 m h}^0$
 - Fitting electroweak data
 - Search for neutral Higgs bosons (h and A)
- $BR(b \rightarrow s\gamma)$ measurements at B-factories
- Anomalous magnetic moment of the muon (BNL)
- Satellite expts WMAP and Planck: density of dark matter (DM), indirect
- Direct searches for DM with ν -detectors and AMS2

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Natural SUSY?

- Light (~ 1 TeV) SUSY particles help to eliminate the hierarchy problem and keep the lightest Higgs-boson light ($\tilde{h}_{MSSM}^0 = H_{SM}$).
- Heavy SUSY particles add huge 2nd-order (log) corrections, ruining the hierarchy elimination.
- Unfortunately, naturalness is less and less probable as the lower limits on SUSY masses grow.

E.g., $ilde{{
m t}}$ is assumed to be the lightest squark, now having $m(ilde{{
m t}})\gtrsim 1$ TeV, and a possible decay of the gluino is to ${
m t} ilde{{
m t}}$...

CMS SUSY summary plot, 2017

Selected CMS SUSY Results* - SMS Interpretation ICHEP '16 - Moriond '17 PP - 0.0 - 49) SUS-16-014 SUS-16-033 OUNHT PP - 0 0.0 - 49) SUS-18-018 SUS-TS-036 00MT2 SUS-16-014 (US-16-03.3 ORMHT) m -88.8 - 16 SUS-16-018 SUS-TG-038 00MTD m -00.0 - 66 P -88.8 - 56 848-18-018 00 848-18-014 9L15-16-033 0EMHTS PP -8 8.8 -11 SUS-16-015 SUS-16-036 06MTD) PP -0 0.0 - 11 SUS-10-016 OF PP -0 9.9 - 11 Gluino SUB-18-019 SUB-16-042 18 h m) PP 0 0.0 - tt PP-0 0.0 -tt : 858-10-020 SEE-16-035 2 same-sign SUS-16-022 SUS-16-04 T Multilepton PP-09.0-tt) PP -0 9.0 - 11 1205-18-030 0 PP-00.0 - tt 00.0 -tt -tej Las = 20 Cin V PP - 0 9. 0 - bt 001 - 00 W US-16-019 SUS-16-042 18 1 of SUS-16-020 SUS-16-03.6 2 same-sign 1.0.5 PP - 0 0 . 0 - 001 - 00 W PP - 0 9 . 9 - 991 - 99 W 5U5-16-020 8U5-16-035 2 seme-sign 66.6 - mi-' 1.) -- 44 (WED) SUS-16-014 (CUS-16-033 00MHT) z=0.5 -99(WD) 5U5-16-022 0015-16-041 Multileo to -- 481 X will f p -H .I -t SUS-16-014 SUS-16-03.3 OFMIN p -tt,t-t SUS-16-018 (EIS-16-03.6 00MT2) p-it,t-t SU8-16-016 OF p-tt,t-t SU8-16-027 SUS-17-00121 opposite-sign p-it i -t i 508-16-028 508-16-05111 m-tt,t-t i SU8-16-029 SUS-16-04.8 OF m-tt ,t -t i 808-16-030 0 pp-itt, t - e z Max exclusion for M_{states} - M _{Lep} < 80 GeV) PP-11.1 → 0 010012 Max exclusion for M_{max} - M up < 80 GeV) **CMS** Preliminary pp-itt,t -e occius ion for M Max exclusion for Massa - M up < 80 GeV) pp-it,t-bff (+body SUS-16-025 SUS-16-04 II 22 and pp→tt,t→bff (4body pp→tt,t→bff (4body Max exclusion for M_____ • M up < 80 GeV) Max exclusion for M____ • M up < 80 GeV) SU6-16-029 SU8-16-0411 0 √s = 13TeV \$538-16-031 11 april pp tt,t -i b-bW 508-16-026 505-16-05111 X 10.5 SUS-16-029 SUS-16-04.9 0 1.00 pp -tt ,t -1; b-- bW' 2-0.5 2-0.5 pp-tt.t $L = 12.9 \text{ fb}^{-1} L = 35.9 \text{ fb}^{-1}$ 2 rep per to-align pp -bb, b + b SUB-10-014 01-10-015 DOMHT pp -- bb, b -- b \$48-16-018 SUS-16-03.0 OFMT2 pp -bb, b -b SUB-10-016 OF pp -bb, b +b q +q,(u, de,s) PP 44.4 +4 : SUS-16-014 SUS-16-03 3 00MHT q +q (u de.s) PP 44.4 +42 \$48-16-018 \$ES-TE-036 OFMT2 1.0 1 1 - 1 - 1 I Sec. 60 -101 $\mathcal{D}_{i,j}$ 2001 2 x=0.5 $\begin{array}{c} p_{0} \rightarrow p_{1} & p_{2} \rightarrow w_{2} & p_{3} \\ p_{0} \rightarrow p_{1} & p_{2} \rightarrow w_{2} & p_{3} \\ p_{0} \rightarrow p_{2} & p_{3} \rightarrow w_{1} & p_{3} \\ p_{0} \rightarrow p_{1} & p_{2} \rightarrow w_{2} & p_{3} \end{array}$ For decays with intermediate mass, mintermediate = x · mMother + (1-x) · mLsp Mass exclusion for M_{mass} - M _{up} < 40 GeV) 2000 200 400 600 800 1000 1200 1400 1600 1800 Mass Scale [GeV] *Observed limits at 95% C.L. - theory uncertainties not included Only a selection of available mass limits. Probe *up to* the quoted mass limit for m =0 GeV unless stated otherwise

Simplified Model Spectrum (SMS) topologies

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CMS limits, 2020: gluino pairs



t the LHC BSM School, ELTE, Budapest, 2021.02.02.

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ATLAS SUSY summary plot, 2017



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CMS: search for exotica



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Search for charged Higgs bosons



M. Aaboud *et al.* [ATLAS Collaboration], Search for charged Higgs bosons produced in association with a top quark and decaying via $H^{\pm} \rightarrow \tau \nu$ using pp collision data recorded at $\sqrt{s} = 13$ TeV by the ATLAS detector, Phys. Lett. B 759 (2016) 555 A. M. Sirunyan *et al.* [CMS Collaboration], Search for charged Higgs bosons produced in vector boson fusion processes and decaying into a pair of W and Z bosons using proton-proton collisions at sqrt(s) = 13 TeV, arXiv:1705.02942 [hep-ex].

BSM School, ELTE, Budapest, 2021.02.02.

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b \rightarrow **s** γ **: FCNC decay**

Flavour-changing neutral current, forbidden in SM

Pure loop effect New physics could affect CLEO, BELLE, BABAR:

 $rac{\Gamma(\mathrm{b}{
ightarrow}\mathrm{s}\gamma)}{\Gamma(\mathrm{b}{
ightarrow}\mathrm{Xe}
u)}=$

$$egin{pmatrix} +0,50\ -0,44 \end{pmatrix} imes 10^{-3} \end{cases}$$

SM: $(3,22\pm0,09) imes10^{-3}$

Agrees with SM ↓ limits MSSM

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CLEO, Cornell



BELLE, Tsukuba

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BABAR, Stanford

CMS H $\rightarrow \gamma \gamma$ in *The Big Bang Theory*



It would be nice if in the (not too distant) future we could see an observed SUSY event in a popular TV show!

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Conclusion

- We have observed the Standard Model Higgs boson or (unfortunately, much less probably) a Higgs boson of a more general model.
- All measured properties are consistent with the predictions for the SM Higgs-boson with a mass of 125 GeV.
- Let us hope for some deviation from the Standard Model (although none seen yet).
- The simplest SUSY models do not seem to be supported by experimental data (g-2, LEP, WMAP, LHC, ...)
- We are looking for and hoping to find new physics (SUSY, Dark Matter, extra Higgs bosons, ...) at the LHC.

Thanks for your attention!

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